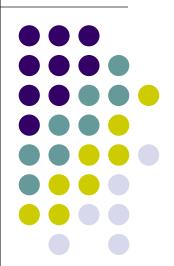
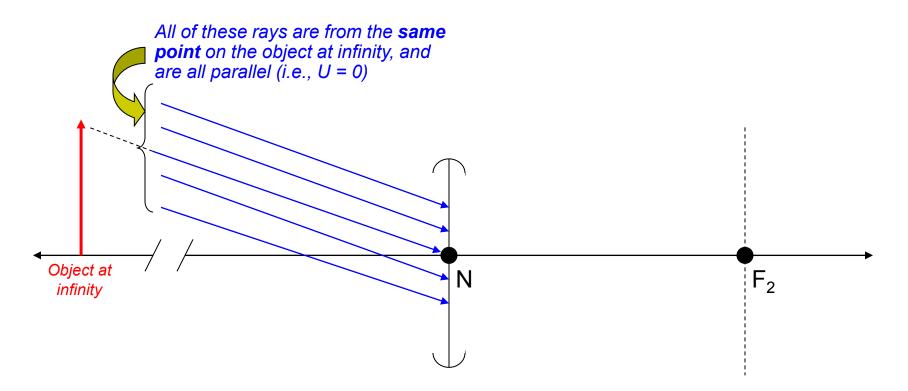
Basic Optics, Chapter 23



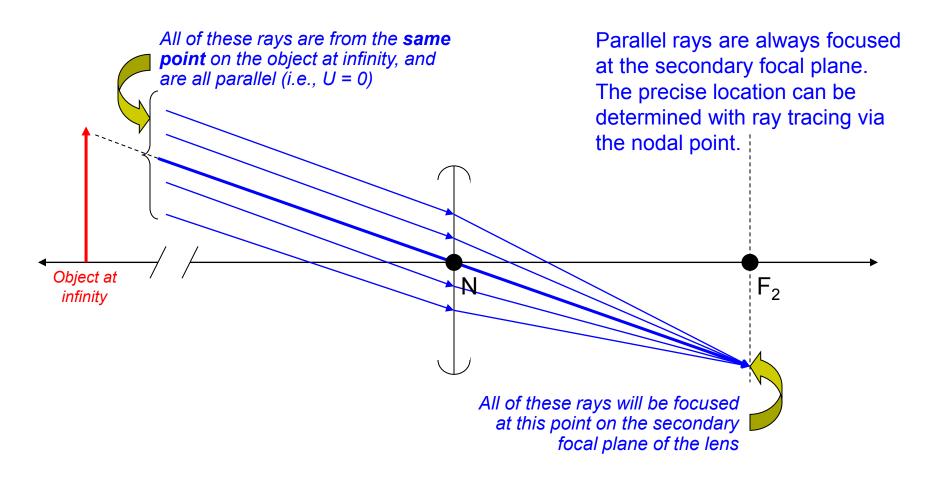


Consider an optical system consisting of a single plus lens and an object at infinity: Where will the rays from the tip of the object be focused?



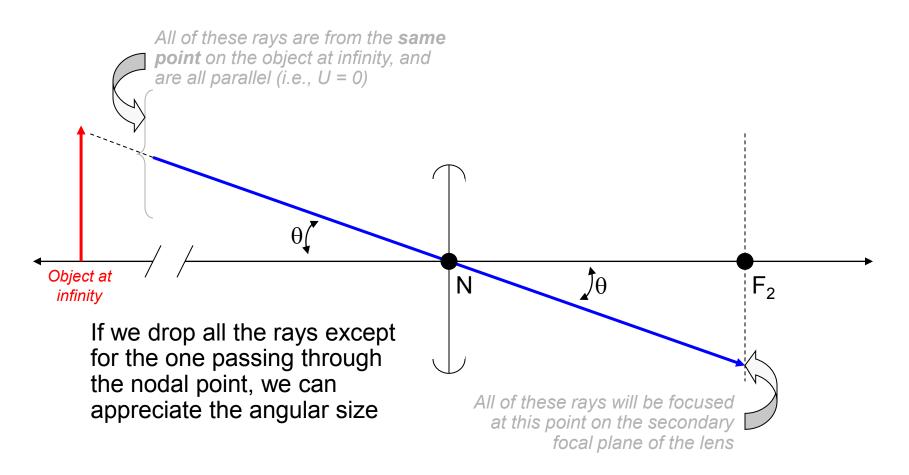


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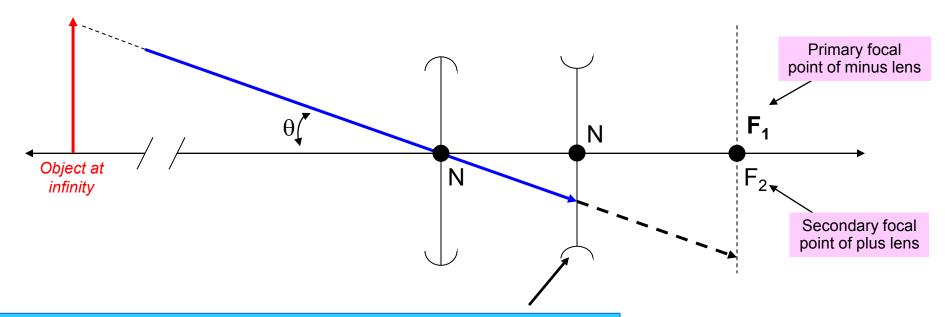


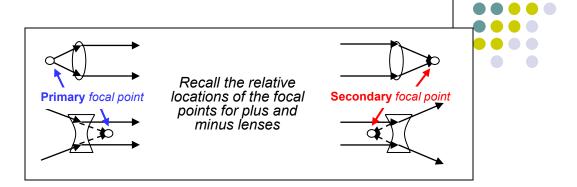


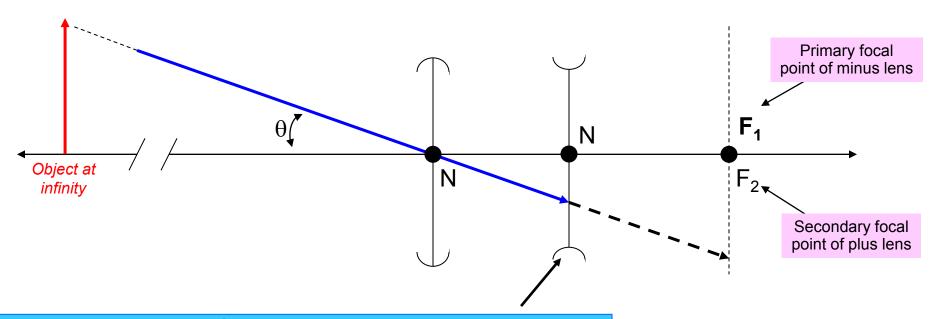
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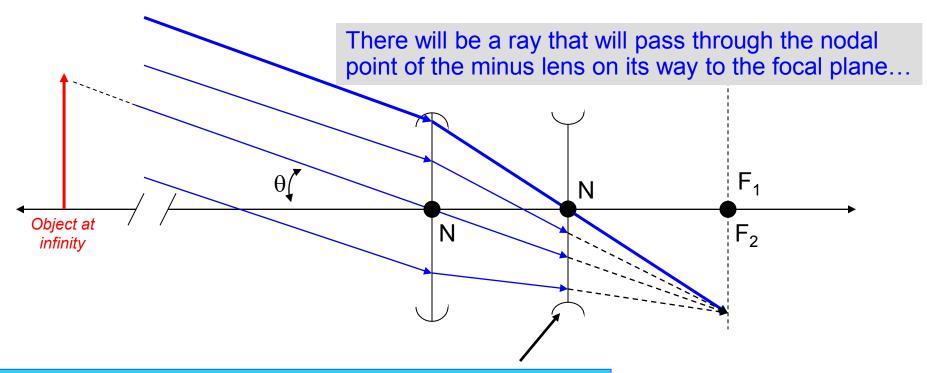




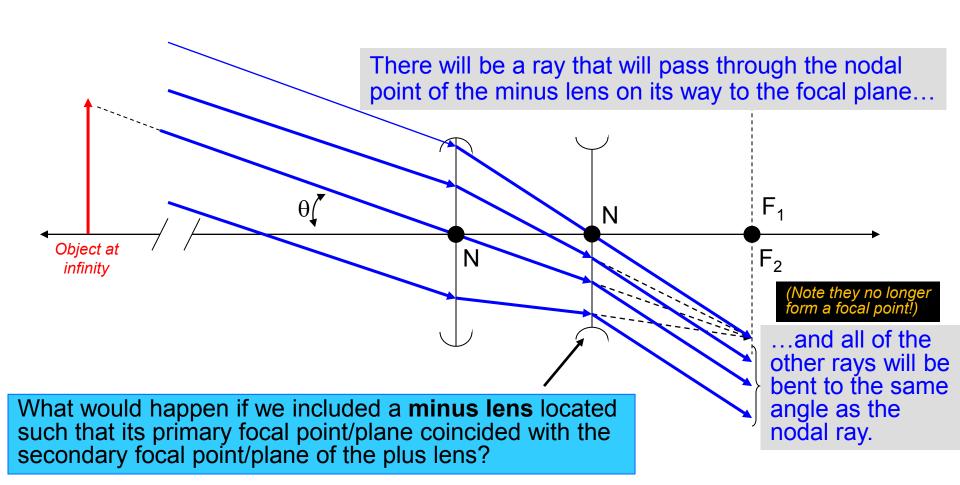




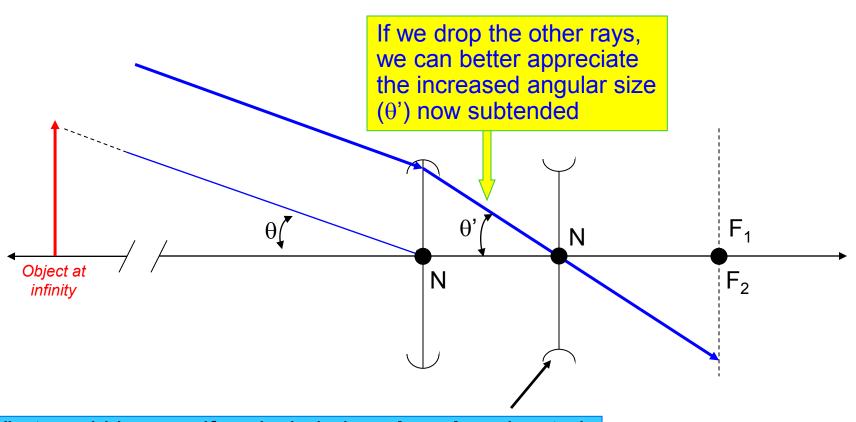


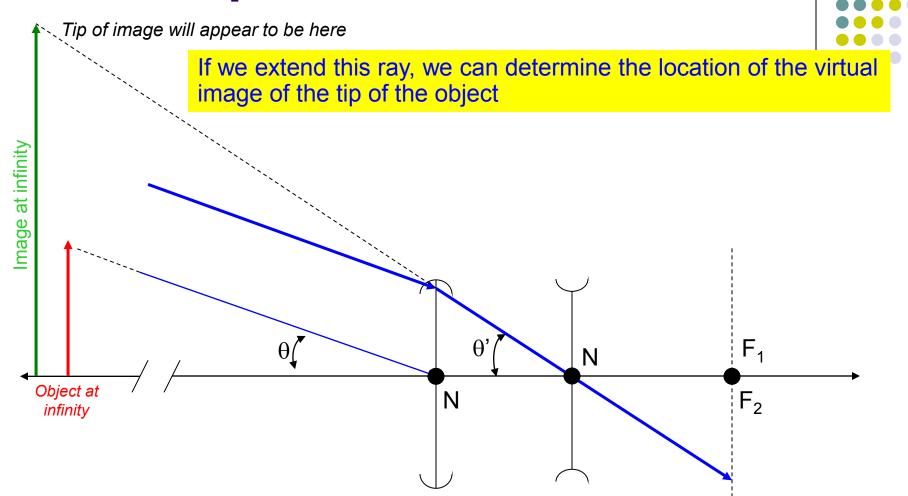










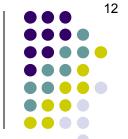


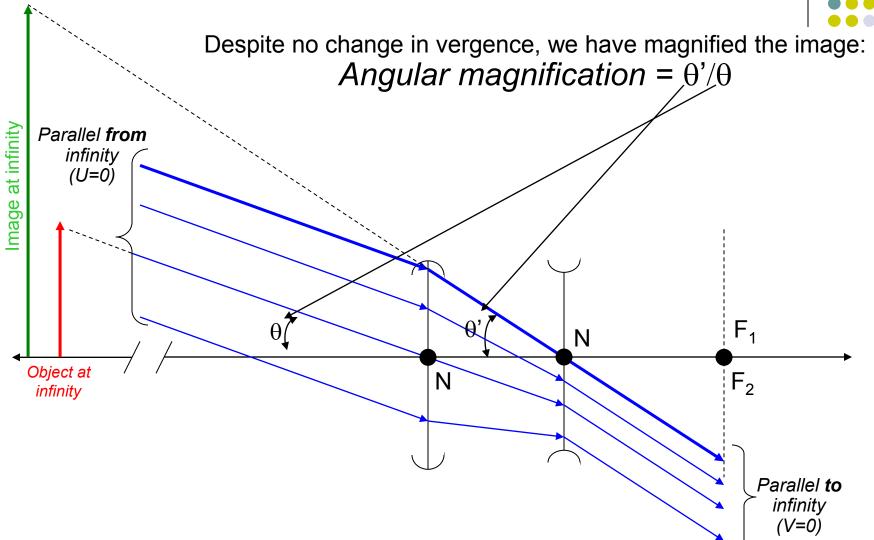
10

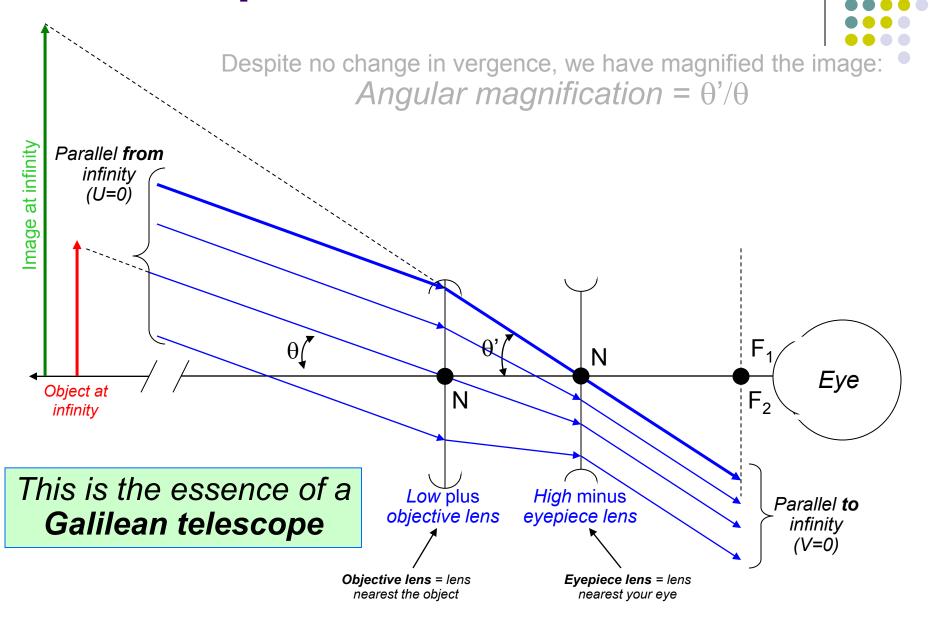
Telescopes Note that the system remains *afocal* (ie, parallel rays entering and leaving) Image at infinity Parallel from infinity (U=0)Ν Object at F_2 N infinity

11

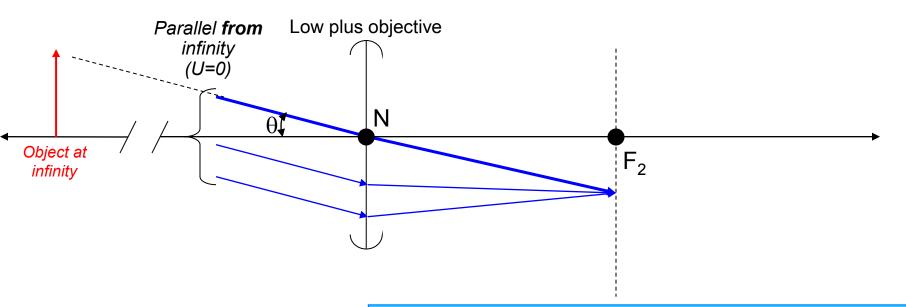
Parallel **to** infinity (V=0)





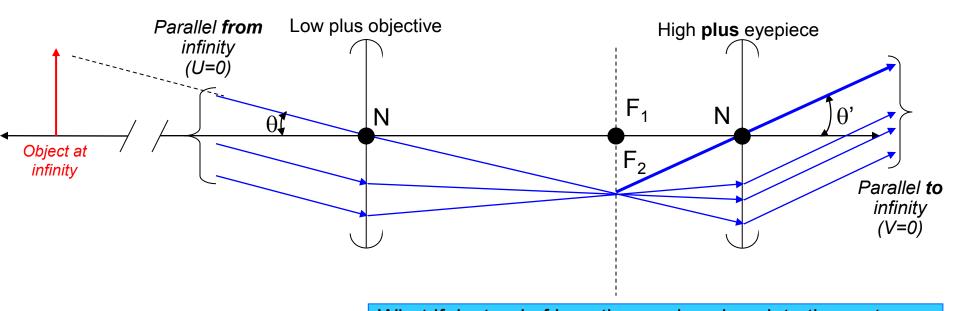




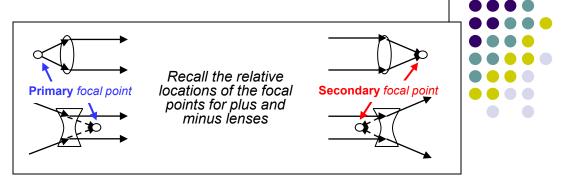


What if, instead of inserting a minus lens into the system...

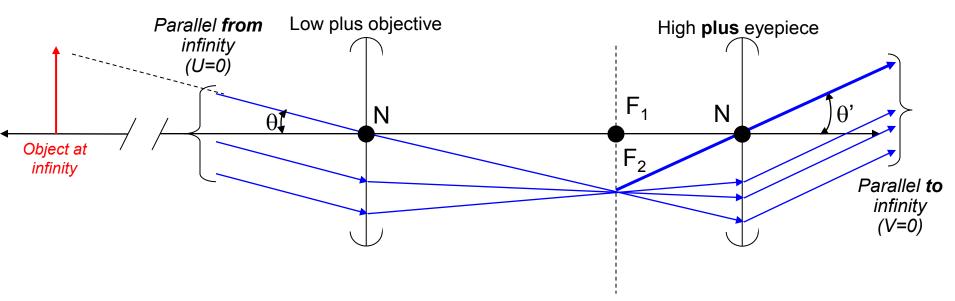




What if, instead of inserting a minus lens into the system... we inserted a **plus lens**, located so its primary focal point coincides with the secondary focal point of the plus lens?

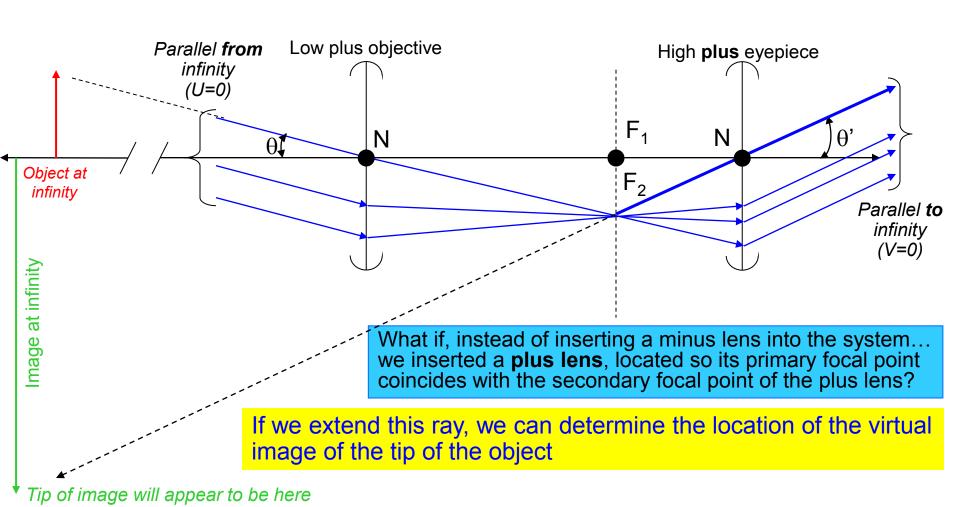


16

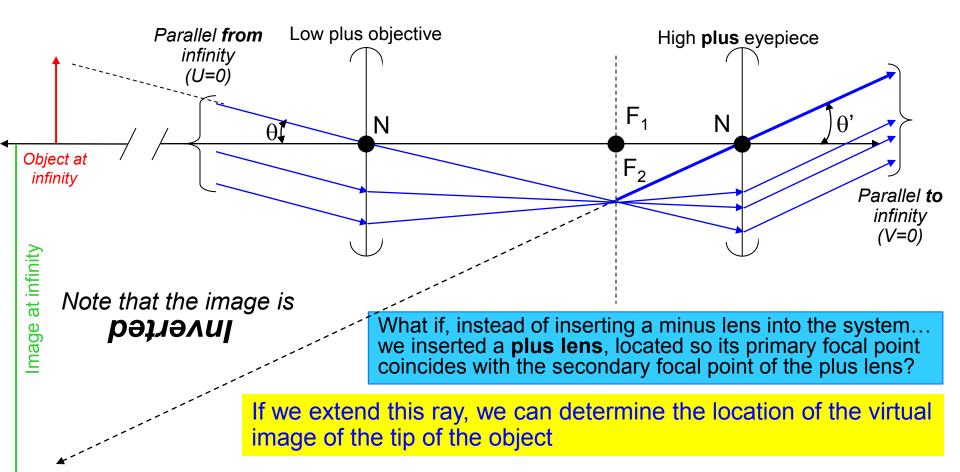


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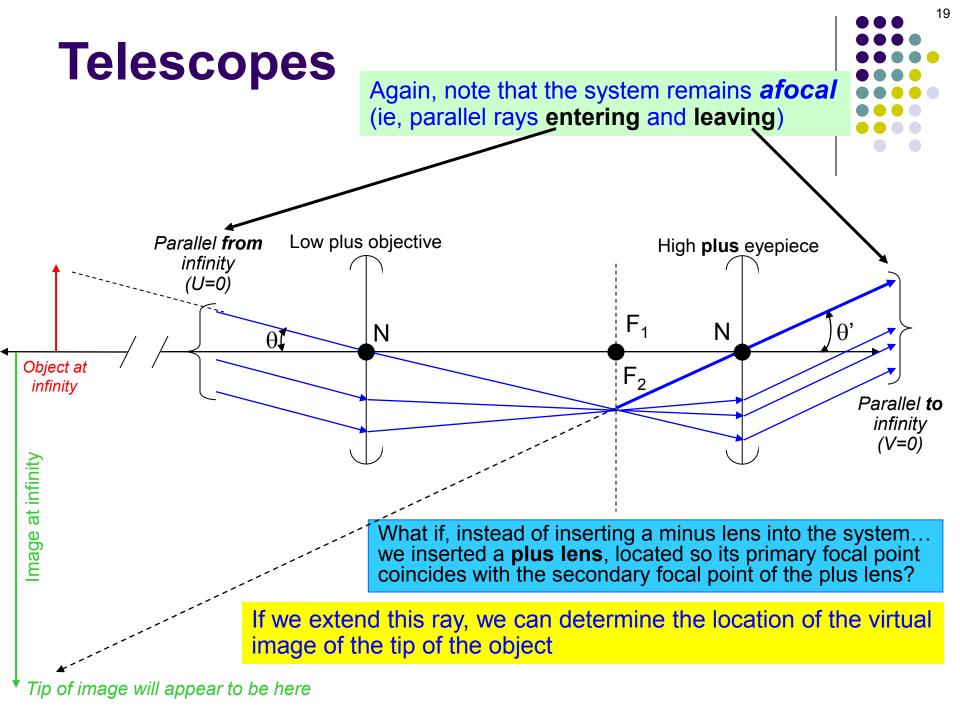






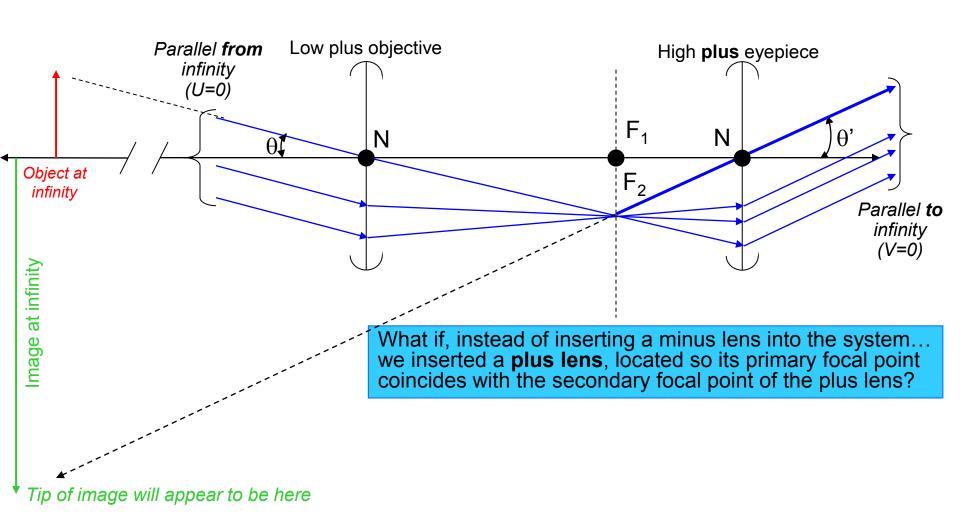


Tip of image will appear to be here



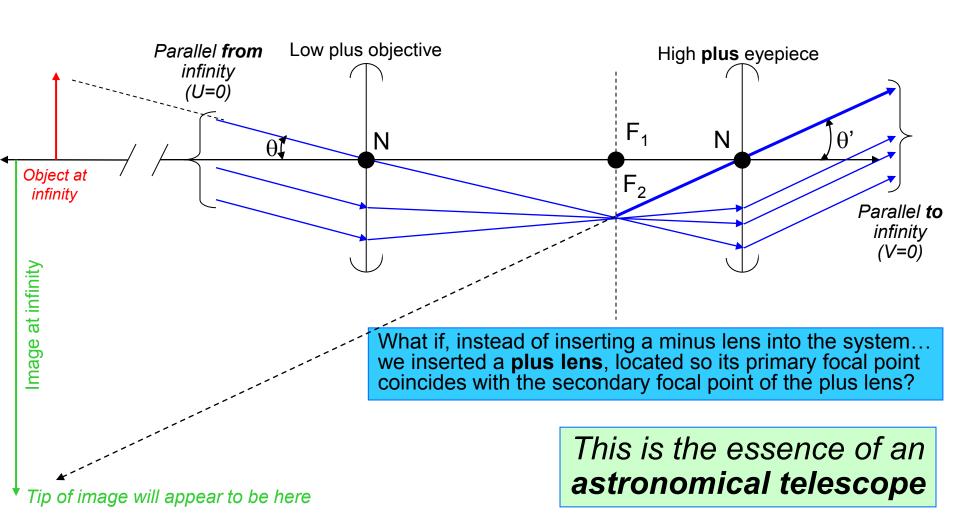
20

Despite no change in vergence, we have magnified the image: $Angular\ magnification = \theta'/\theta$





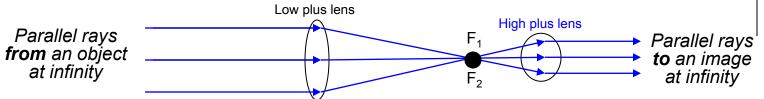
Despite no change in vergence, we have magnified the image: $Angular\ magnification = \theta'/\theta$





To reiterate: Telescopes come in two basic flavors—those with a high **plus** eyepiece lens, and those with a high **minus** eyepiece lens.

Astronomical (Keplerian) telescope

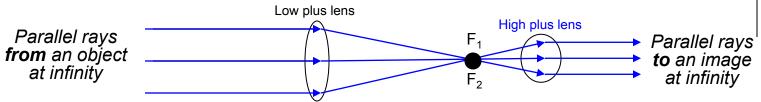


To reiterate: Telescopes come in two basic flavors—those with a high **plus** eyepiece lens, and those with a high **minus** eyepiece lens.

High-plus-eyepiece telescopes are called *astronomical* (or *Keplerian*) telescopes;



Astronomical (Keplerian) telescope

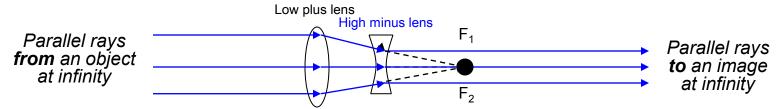


To reiterate: Telescopes come in two basic flavors—those with a high **plus** eyepiece lens, and those with a high **minus** eyepiece lens.

High-plus-eyepiece telescopes are called *astronomical* (or *Keplerian*) telescopes; and

high-minus-eyepiece telescopes are called *Galilean* (or *terrestrial*) telescopes.

Galilean (terrestrial) telescope





Angular magnification = θ'/θ is difficult to work with.



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Fortunately, for reasonably small angles, this can be well approximated by:



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28

Telescopes

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Don't forget this minus sign! It keeps the magnification value consistent with our image orientation sign convention

29

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$$\frac{Angular}{mag} = -\frac{\text{Eyepiece lens}}{\text{Objective lens}} = -\frac{\text{Plus}}{\text{Plus}} = (-)$$

Astronomical telescope (image is **pəɪɹəʌui**)

30

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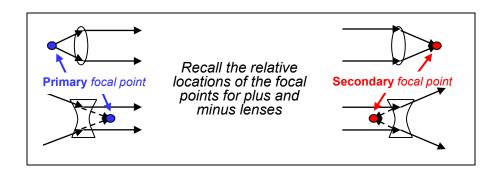
$$\frac{Angular}{mag} = \frac{\text{Eyepiece lens}}{\text{Objective lens}} = \frac{\text{Minus}}{\text{Plus}} = (+)$$

Astronomical telescope (image is **pəɪɹəʌui**)

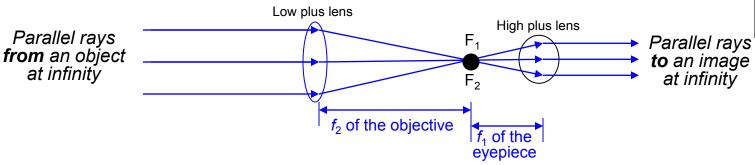
Galilean telescope (image is *upright*)



For a telescope to function, the primary focal point of the eyepiece must overlap the secondary focal point of the objective. This determines the separation between the two lenses.



Astronomical (Keplerian) telescope

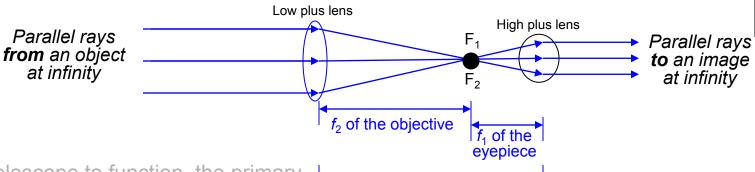


For a telescope to function, the primary focal point of the eyepiece must overlap the secondary focal point of the objective. This determines the separation between the two lenses.



Astronomical (Keplerian) telescope

Separation = $f_2 + f_1$



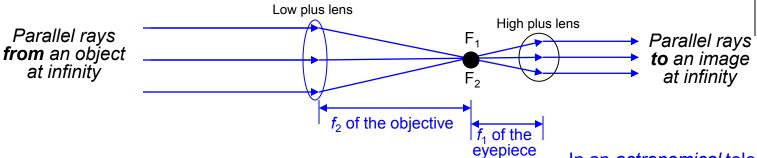
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For a telescope to function, the primary focal point of the eyepiece must overlap the secondary focal point of the objective.

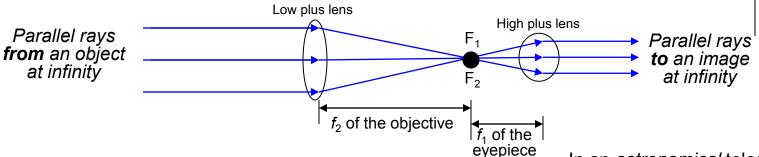
This determines the separation between the two lenses.

In an *astronomical* telescope, the separation is equal to the **sum** of the focal lengths.

34

Astronomical (Keplerian) telescope

Separation = $f_2 + f_1$



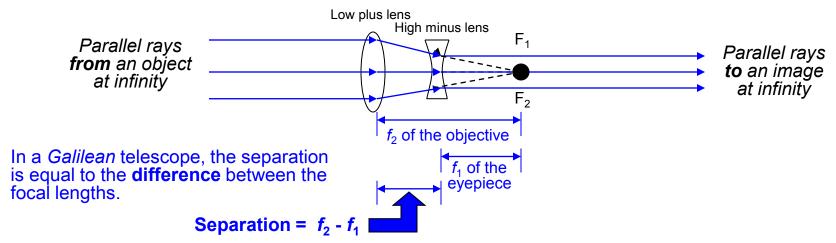
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35

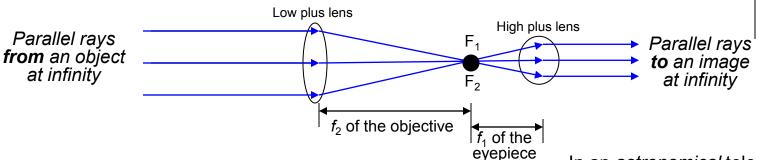
This determines the separation between the two lenses.

Galilean (terrestrial) telescope



Astronomical (Keplerian) telescope

Separation = $f_2 + f_1$



For a telescope to function, the primary focal point of the eyepiece must overlap the secondary focal point of the objective.

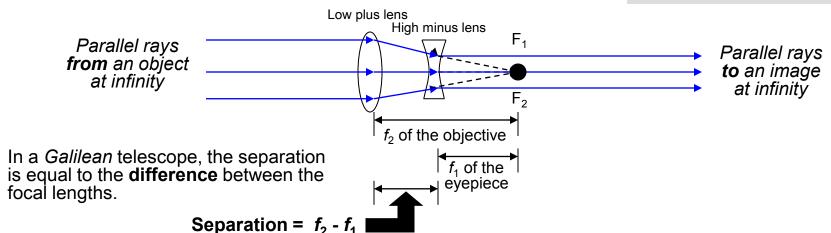
This determines the separation between the two lenses.

In an *astronomical* telescope, the separation is equal to the **sum** of the focal lengths.

36

For this and other reasons, Galilean scopes tend to be **smaller** and **lighter** than astronomical scopes.

Galilean (terrestrial) telescope



Astronomical (Keplerian) telescope

High plus lens

Parallel rays from an object at infinity

For a telescope to function,

focal point of the eyepiece i

the secondary focal point of

The main other reason being that astronomical telescopes require *prisms* to flip the inverted image into the upright position, which add considerably to their size and weight!

Low plus lens

This determines the separation between the two lenses.

In an astronomical telescope, the separation is equal to the **sum** of the focal lengths.

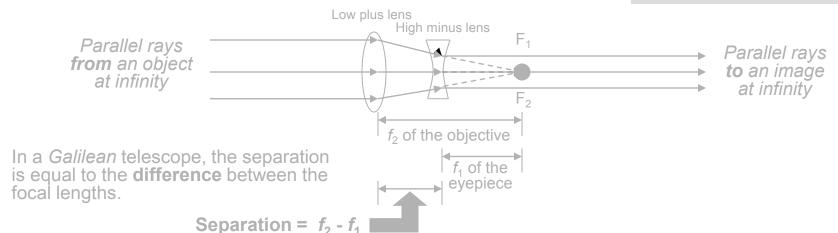
Parallel rays

to an image

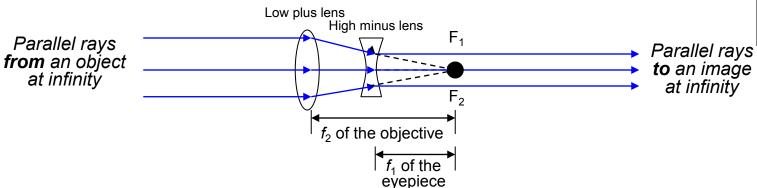
at infinity

For this and other reasons, Galilean scopes tend to be smaller and lighter than astronomical scopes.

Galilean (terrestrial) telescope

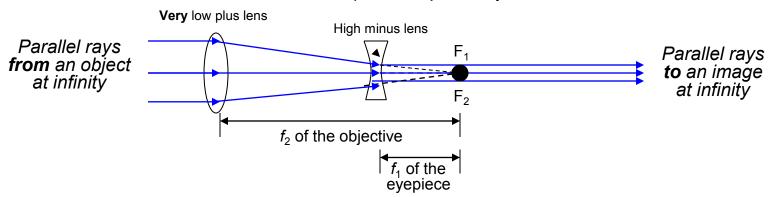


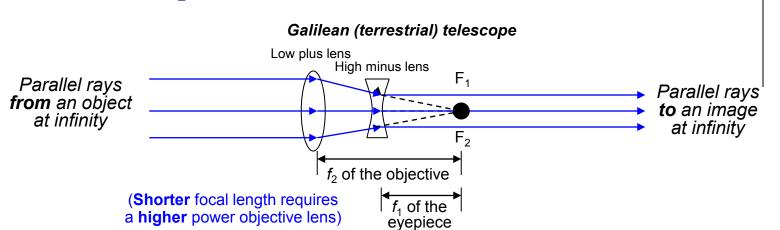
Galilean (terrestrial) telescope



Let's compare two Galilean telescopes. Note the difference in focal lengths.

Galilean (terrestrial) telescope





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The greater lens separation below necessitates a **decrease** in power for the objective lens (and vice versa). Will this increase, decrease or leave unaffected the overall power of the telescope?

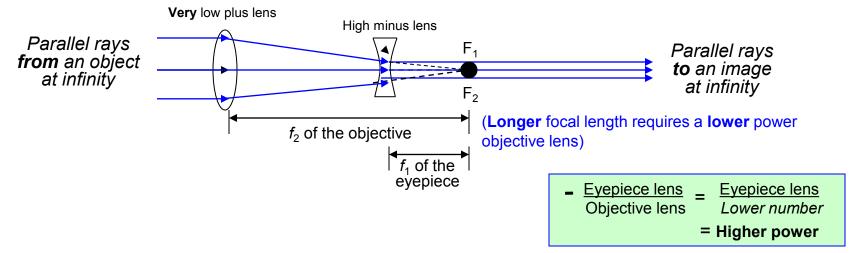
Galilean (terrestrial) telescope

a **lower** power objective lens)

Galilean (terrestrial) telescope Low plus lens High minus lens $\mathsf{F}_{\scriptscriptstyle{1}}$ Parallel rays Parallel rays from an object to an image at infinity at infinity (Shorter focal length requires a higher power f_2 of the objective objective lens) eyepiece Eyepiece lens _ Eyepiece lens Objective lens Higher number = Lower power **Increase** overall power.

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Galilean (terrestrial) telescope





	Field of view	Maximum practical magnification		
Galilean				
Astronomical				



	Field of view	Bolativa Siza	Relative
Galilean	Smaller		
Astronomical	Larger		



	Field of view	Maximum practical magnification		Relative
Galilean	Smaller			
Astronomical	Larger			



	Field of view	Maximum practical magnification		Relative
Galilean	Smaller	4x		
Astronomical	Larger	10x		



	Field of view	Maximum practical magnification	Relative size	
Galilean	Smaller	4x		
Astronomical	Larger	10x		



	Field of view	Maximum practical magnification	Relative size	lmage orientation	
Galilean	Smaller	4x	Small/light		
Astronomical	Larger	10x	Large/heavy		



	Field of view	Maximum practical magnification	Relative size	Image orientation	
Galilean	Smaller	4x	Small/light		
Astronomical	Larger	10x	Large/heavy		



	Field of view	Maximum practical magnification	Relative size	Image orientation	Relative
Galilean	Smaller	4x	Small/light	Upright	
Astronomical	Larger	10x	Large/heavy	Inverted (unless prisms used)	



	Field of view	Maximum practical magnification	Relative size	Image orientation	Relative cost
Galilean	Smaller	4x	Small/light	Upright	
Astronomical	Larger	10x	Large/heavy	Inverted (unless prisms used)	



	Field of view	Maximum practical magnification	Relative size	Image orientation	Relative cost
Galilean	Smaller	4x	Small/light	Upright	Cheap
Astronomical	Larger	10x	Large/heavy	Inverted (unless prisms used)	Expensive



At this juncture, you should assess your Optics knowledge by taking Quiz 5 (slide-set BO31). After that, resume the tutorial with slide-set BO24.